

**BROCK UNIVERSITY**

Test 1: June 2017

Course: ASTR 1P02

Examination date: 17 June 2017

Time of Examination: 13:00 – 13:50

Number of pages: 10

Number of students: 366

Time limit: 50 min

Instructor: S. D'Agostino

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**Answer all questions on the scantron sheet provided. No aids permitted except for a non-programmable calculator. Each question is worth 1 mark. Total number of marks: 50.**

**DO NOT WRITE YOUR ANSWERS ON YOUR QUESTION PAGE. DOING SO WILL RESULT IN AN ASSIGNED GRADE OF ZERO.**

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1. The cool, early phase of a star is called a
  - (a) pre-star.
  - (b) juvenile star.
  - (c) protostar.
  - (d) proton star.
  
2. Main-sequence stars are in hydrostatic equilibrium, which means the inward gravitational forces are balanced by outward forces due to
  - (a) magnetic flux.
  - (b) hydrostatic fusion.
  - (c) neutrino convection.
  - (d) gas pressure.
  
3. For a star in hydrostatic equilibrium, an important way in which energy is transported in the star's interior is
  - (a) nuclear fission.
  - (b) nuclear fusion.
  - (c) conduction.
  - (d) fluorescence.
  - (e) [None of the above.]
  
4. In the Sun, the convective zone is
  - (a) near the surface of the Sun.
  - (b) near the core of the Sun.
  - (c) about mid-way between the core and the surface of the Sun.
  - (d) throughout the entire Sun.
  - (e) [The Sun has no convective zone.]

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5. The processes that produce radiant energy in the Sun can be summarized as \_\_\_\_\_ hydrogen nuclei combine to produce a helium nucleus and energy is released.
- (a) 2
  - (b) 3
  - (c) 4
  - (d) 5
  - (e) [None of the above.]
6. As a contracting protostar heats up, if the core temperature reaches \_\_\_\_\_ then fusion of hydrogen into helium begins and the protostar becomes a main-sequence star.
- (a) 10 thousand degrees K
  - (b) 10 million degrees K
  - (c) 10 billion degrees K
  - (d) 10 trillion degrees K
7. High temperatures and pressures are required to produce nuclear reactions because this helps
- (a) convince atomic nuclei to get to know each other.
  - (b) atomic nuclei attract each other.
  - (c) to overcome the electrical repulsion between atomic nuclei.
  - (d) to overcome the gravitational attraction between atomic nuclei.
8. A vast number of photinos is produced in the Sun's
- (a) core.
  - (b) atmosphere.
  - (c) convective zone.
  - (d) conductive zone.
  - (e) [The Sun does NOT produce photinos.]
9. As a main-sequence star evolves, steadily converting hydrogen to helium in its core, the core of the star gradually
- (a) becomes cooler and contracts.
  - (b) becomes cooler and expands.
  - (c) becomes hotter and expands.
  - (d) becomes hotter and contracts.
  - (e) [There is no change in the core temperature and size.]

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10. As a main-sequence star evolves, steadily converting hydrogen to helium in its core, the outer layers of the star gradually
- (a) become cooler and contract.
  - (b) become cooler and expand.
  - (c) become hotter and expand.
  - (d) become hotter and contract.
  - (e) [There is no change in the temperature and size of the outer layers.]
11. The Sun began its life as a
- (a) hot, agitated cloud of gas.
  - (b) white dwarf.
  - (c) red giant.
  - (d) neutron star.
  - (e) [None of the above.]
12. Heat flows from a star's core to its surface because
- (a) the star's core is hotter than its surface.
  - (b) of the star's high core gravity.
  - (c) of the star's high core mass.
  - (d) of the star's high core electron flux.
13. For a low-mass main-sequence star, hydrogen is fused into helium mainly by
- (a) the triple-alpha process.
  - (b) the CNO cycle.
  - (c) the proton-proton chain.
  - (d) the Krebs cycle.
14. For a high-mass main-sequence star, hydrogen is fused into helium mainly by
- (a) the triple-alpha process.
  - (b) the CNO cycle.
  - (c) the proton-proton chain.
  - (d) the Krebs cycle.

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15. Main-sequence stars with masses less than 0.4 solar masses are
- (a) brown dwarfs.
  - (b) red dwarfs.
  - (c) yellow dwarfs.
  - (d) white dwarfs.
16. An important way that astronomers measure distances to other galaxies is using
- (a) planetary nebulae.
  - (b) Cepheid variable stars.
  - (c) Type IIB quasars.
  - (d) the method of galactic halos.
  - (e) the method of giant molecular clouds.
17. The mass of Star A is much greater than the mass of Star B, and they are both main-sequence stars. The rate at which Star A consumes its hydrogen fuel is \_\_\_\_\_ the rate at which Star B consumes its hydrogen fuel.
- (a) greater than
  - (b) about the same as
  - (c) less than
  - (d) [Not enough information is given.]
18. Small, dark clouds called Bok globules are found in and near
- (a) black holes.
  - (b) pulsars.
  - (c) star-forming regions of space.
  - (d) type Ia supernovae.
19. A planetary nebula is a cloud of gas and dust surrounding
- (a) a Jovian planet.
  - (b) a terrestrial planet.
  - (c) an exoplanet.
  - (d) a dwarf planet.
  - (e) [None of the above.]

20. The formation of “heavy elements” (nucleosynthesis) occurs primarily during
- (a) the collapse of galactic nuclei after a collision.
  - (b) the core collapse of a high-mass star.
  - (c) the collapse of a T-Tauri star as it reaches the main sequence.
  - (d) the collapse of a planetary nebula.
21. The distances to the most distant galaxies are measured using
- (a) Type Ia supernovae.
  - (b) Type II supernovae.
  - (c) Cepheid variables.
  - (d) T Tauri stars.
22. Red dwarfs have been observed to finish consuming all of their available hydrogen fuel
- (a) about once per month.
  - (b) about once per decade.
  - (c) about once per century.
  - (d) about once per thousand years.
  - (e) [This has never been observed.]
23. Degenerate matter is found
- (a) in the atmosphere of a low-mass main-sequence star.
  - (b) in most cool, low-density, giant molecular clouds.
  - (c) in the core of a white dwarf.
  - (d) in the spectrum of a type Ia supernova.
  - (e) only in laboratories on Earth.
24. An accretion disk is formed from
- (a) the core collapse of a red giant.
  - (b) gas flowing from a companion star towards a white dwarf.
  - (c) clumping in a planetary nebula.
  - (d) clumping in a giant molecular cloud.
25. Most of the energy produced in Type II supernovae explosions is carried away by
- (a) neutrons.
  - (b) protons.
  - (c) radio waves.
  - (d) visible light.
  - (e) [None of the others.]

26. The surface temperature of a white dwarf is typically
- (a) less than the Sun's surface temperature.
  - (b) about the same as the Sun's surface temperature.
  - (c) greater than the Sun's surface temperature.
27. The density of a white dwarf is typically
- (a) about the same as the Sun's density.
  - (b) about the same as the Earth's density.
  - (c) greater than a neutron star's density.
  - (d) greater than the Sun's density.
  - (e) [None of the others.]
28. In a white dwarf, equilibrium is maintained by a balance of
- (a) core temperature and core mass.
  - (b) core temperature and core pressure.
  - (c) electron degeneracy pressure and gravity.
  - (d) electronegativity pressure and electron degeneracy pressure.
  - (e) electronegativity pressure and gravity.
29. If the mass of a neutron star were to suddenly increase to greater than about \_\_\_\_\_, then it would collapse into a black hole.
- (a) 3 solar masses
  - (b) 30 solar masses
  - (c) 300 solar masses
  - (d) 3,000 solar masses
30. When a type Ia supernova explodes, it leaves behind
- (a) a white dwarf as a central remnant.
  - (b) a neutron star as a central remnant.
  - (c) a black hole as a central remnant.
  - (d) no central remnant.
  - (e) [Either (b) or (c) could occur.]

31. When a type II supernova explodes, it leaves behind
- (a) a white dwarf as a central remnant.
  - (b) a neutron star as a central remnant.
  - (c) a black hole as a central remnant.
  - (d) no central remnant.
  - (e) [Either (b) or (c) could occur.]
32. Neutron stars typically
- (a) have smaller diameters and are less dense than white dwarfs.
  - (b) have smaller diameters and are more dense than white dwarfs.
  - (c) have larger diameters and are less dense than white dwarfs.
  - (d) have larger diameters and are more dense than white dwarfs.
33. Type Ia supernovae
- (a) have hydrogen spectral lines in their spectra.
  - (b) do not have hydrogen spectral lines in their spectra.
34. Type II supernovae
- (a) have hydrogen spectral lines in their spectra.
  - (b) do not have hydrogen spectral lines in their spectra.
35. Binary pulsars emit \_\_\_\_\_ that can be used to test \_\_\_\_\_ .
- (a) electromagnetic waves / Planck's theory of black body radiation
  - (b) gravitational waves / Einstein's theory of general relativity
  - (c) magnetohydrodynamic waves / Alfven's theory of plasma oscillations
  - (d) shock waves / Mach's theory of ultrasonic oscillations
36. In an X-ray binary star, the object emitting the X-rays is
- (a) a main-sequence star.
  - (b) a neutron star.
  - (c) a red dwarf.
  - (d) a white dwarf.
37. As the 20th century began, astronomers generally believed that the Sun was located near the centre of a wheel-shaped star system that they estimated had a diameter of
- (a) about 15 light-years.
  - (b) about 15,000 light-years.
  - (c) about 15,000,000 light-years.
  - (d) about 15,000,000,000 light-years.

38. That the Milky Way is made up of an enormous number of individual stars was confirmed by Galileo in 1609, but was first suggested by
- (a) Xander Bogaerts in 200 BCE.
  - (b) Pliny the Elder in 1185.
  - (c) Tycho Brahe in 1583.
  - (d) William and Caroline Herschel in 1785.
  - (e) [None of the above.]
39. The first reasonably good approximation for the size of the Milky Way was determined by
- (a) Svetožar Gligorić.
  - (b) Stephen Hawking.
  - (c) Theodore Roosevelt.
  - (d) Vera Rubin.
  - (e) Harlow Shapley.
40. The first reasonably good approximation for the size of the Milky Way was determined by measuring the distances to
- (a) giant molecular clouds.
  - (b) globular star clusters.
  - (c) pulsars.
  - (d) comets.
41. Early 20th-century astronomers observed that the Milky Way's globular star clusters appeared to be centred at a point many thousands of light-years away in the direction of the constellation
- (a) Aries.
  - (b) Gemini.
  - (c) Leo.
  - (d) Sagittarius.
42. One way astronomers deduce the mass of the Milky Way is to study
- (a) the orbital motions of stars in the galaxy.
  - (b) the motions of stars within globular clusters.
  - (c) the periods of eclipsing binary pulsars.
  - (d) the spectra of white dwarfs in different sectors of the galaxy.

43. The number of stars in the Milky Way is approximately
- (a) 100 thousand.
  - (b) 100 million.
  - (c) 100 billion.
  - (d) 100 trillion.
44. Density wave theory is currently the most popular accepted explanation for the formation and evolution of the Milky Way's
- (a) central bulge.
  - (b) halo.
  - (c) dark matter.
  - (d) spiral arms.
45. Population I stars are typically found in the Milky Way's
- (a) halo and bulge, have eccentric orbits, and have relatively low heavy-element content.
  - (b) halo and bulge, have eccentric orbits, and have relatively high heavy-element content.
  - (c) spiral arms, have approximately circular orbits, and have relatively high heavy-element content.
  - (d) spiral arms, have approximately circular orbits, and have relatively low heavy-element content.
46. Population II stars are typically found in the Milky Way's
- (a) halo and bulge, have eccentric orbits, and have relatively low heavy-element content.
  - (b) halo and bulge, have eccentric orbits, and have relatively high heavy-element content.
  - (c) spiral arms, have approximately circular orbits, and have relatively high heavy-element content.
  - (d) spiral arms, have approximately circular orbits, and have relatively low heavy-element content.
47. The Milky Way rotates in such a way that
- (a) the rotation rate is nearly the same at all points of the galaxy.
  - (b) the central parts of the galaxy rotate faster than the outer parts.
  - (c) the central parts of the galaxy rotate slower than the outer parts.
  - (d) [The Milky Way does not rotate.]

48. The Milky Way's disk contains most of its
- (a) younger stars.
  - (b) middle-aged stars.
  - (c) older stars.
  - (d) dead stars.
49. According to current theories of the Milky Way's evolution, the shape of the Milky way is predicted to become \_\_\_\_\_ as it evolves.
- (a) flatter.
  - (b) rounder.
  - (c) more twisty.
  - (d) separated into different parts.
  - (e) [None of the above.]
50. There is so much interstellar dust in the Milky Way's disk that
- (a) inhabitants of the Ophiuchus sector suffer from respiratory problems.
  - (b) the nucleus of the Milky Way can't be observed at radio wavelengths.
  - (c) the nucleus of the Milky Way can't be observed at visual wavelengths.
  - (d) the nucleus of the Milky Way can only be observed using neutrinos.